

Physics of Galaxies

ANSWERS: HW SET NUMBER 1

1. a is semi-major axis, b is semi-minor axis, $n=[10\ e]=[10\ (1-b/a)]$, where e is ellipticity and square brackets refer to integer part **[2 marks]**.

For the left figure $n=[10\ e]=[10\ (1-45\text{mm}/45\text{mm})]=0$, thus this is E0

For the middle figure $n=[10\ e]=[10\ (1-13\text{mm}/45\text{mm})]=7$, thus this is E7

For the right figure $n=[10\ e]=[10\ (1-27\text{mm}/45\text{mm})]=4$, thus this is E4 **[3 marks]**

Sa→Sc size of bulge decreases,

SBa→SBc size of bar decreases,

In the both cases winding of spiral arms decreases **[3 marks]**.

2. We have

$$H_o = h \times \frac{100 \times 10^3 (\text{m s}^{-1})}{3.1 \times 10^{22} (\text{m})} = h \times 3.2 \times 10^{-18} \text{s}^{-1}$$

so that

$$\begin{aligned} \frac{1}{H_o} &= \frac{1}{h} \times \frac{10^{18}}{3.2 \times 3.2 \times 10^7} \text{ yr} \\ &\approx \frac{10^{10}}{h} \text{ yr} \end{aligned} \quad \textbf{[2 marks]}$$

Hubble measured h to be 5, giving H_o^{-1} – which is roughly the age of the universe – as 2×10^9 yr, less than the age, for example, of globular clusters in the Galaxy, which is problematic. **[2 marks]**

3. Assuming the observer is in the midplane, the pathlength through the dust for a source at Galactic latitude b is $x = h/\sin(b) = h \csc(b)$. The flux is therefore reduced from F_o (the value without extinction) to

$$F = F_o \exp(-k h \csc b) \quad \textbf{[1 mark]}$$

The apparent magnitude is then changed by

$$m - m_o = -2.5 \log_{10}(F/F_o)$$

$$= - (2.5/\log_e 10) \log_e (F/F_o)$$

$$= 1.09 kh \csc b.$$

[3 marks] [1+3=4 marks]

[Note $\log_a x = \log_b x / \log_b a$]

[Total marks available 16]