

Course Work 6

Q1.

a) A binary system consists of two neutron stars of the same mass M . The orbital period of the system is P . Using Newtonian mechanics, estimate to an order of magnitude the separation between the neutron stars, r , and the fractional relativistic corrections to the orbital motion.

b) Evaluate the relativistic corrections if $P = 8 \text{ min}$ and $M = 1.5M_\odot$. Compare your estimate with relativistic effects in the solar system. It is known that the perihelion shift of Mercury is $43''$ per century. What analogous shift can you expect in the case of the binary system of neutron stars? (Hint: The relativistic shift per one orbital period is of order r_g/r , where r_g is gravitational radius of the neutron star.)

Q2.

The quadrupole formula for the metric perturbation associated with gravitational waves is given by

$$h_{\alpha\beta} = -\frac{2G}{3c^4 R} \frac{d^2 D_{\alpha\beta}}{dt^2} (t - R/c),$$

where R is the distance to the source of the gravitational waves and

$$D_{\alpha\beta} = \int (3x_\alpha x_\beta - r^2 \delta_{\alpha\beta}) dM$$

is the quadrupole tensor of the source. Consider a mass m moving along circular orbit around the black hole of mass M , assuming that $m \ll M$.

a) Show that all the amplitudes $h_{\alpha\beta}$ of gravitational wave, emitted by such system, are periodic functions of time with $\omega = 2\omega_0$, where $\omega_0 = 2\pi/T$, and T is the orbital period.

b) Show that, to an order of magnitude (omitting the indices α and β)

$$h \approx \frac{r_g}{R} \left(\frac{R_g \omega}{c} \right)^{2/3},$$

where r_g is the gravitational radius of the mass m and R_g is the gravitational radius of the black hole.

Q3.

The future LISA mission will be able to detect gravitational waves with $h > 10^{-23}$, if $10^{-4} \text{ Hz} < \omega < 3 \cdot 10^{-3} \text{ Hz}$. From what distance will it be possible to detect gravitational radiation from the binary system, containing the black hole of mass $m = 3M_\odot$, moving along a circular orbit with radius $r = 10^4 R_g$ around the massive black hole of mass $M = 10^3 M_\odot$?