

c. Course work 2

Q1. [25 Marks]

a) Starting with the Newtonian deceleration equation

$$\ddot{R} = -\frac{4\pi G\rho}{3}R, \quad (\text{C.1})$$

derive the Friedman equation in the form

$$\dot{R}^2 = \frac{8\pi G\rho}{3}R^2 - kc^2, \quad (\text{C.2})$$

where ρ is homogeneous density and k is a constant. [10 Marks]

b) Show that the following relationships

$$\Omega(t) = 2q(t), \quad \text{and} \quad kc^2 = R(t)^2 H(t)^2 (\Omega(t) - 1) \quad (\text{C.3})$$

between deceleration

$$q(t) = -\frac{R(t)\ddot{R}(t)}{\dot{R}^2(t)}, \quad (\text{C.4})$$

the dimensionless density

$$\Omega(t) = \frac{8\pi G\rho(t)}{3H^2(t)}, \quad (\text{C.5})$$

the Hubble constant $H(t)$ and the scale factor $R(t)$ are valid for any moment of time (not only at the present moment). [15 Marks]

Q2. [25 Marks]

a) Show that in the Newtonian cosmological model with $k=+1$ the scale factor, $R(t)$, as a function of time can be presented in the following parametric form

$$R = a(1 - \cos \eta), \quad t = \frac{a}{c}(\eta - \sin \eta), \quad (\text{C.6})$$

where a is a constant and the parameter η goes from 0 (the Big Bang) to 2π (the Big Crunch). [15 Marks]

b) Calculate the maximal value of the scale factor R in terms of the main cosmological parameters. [10 Marks]

Q3. [20 Marks]

Find in a parametric form the dependence of the Hubble constant, $H(t)$, on time in Newtonian cosmological model with $k = 1$.

Q4. [30 Marks]

a) Motivate that the fate of the Universe can be predicted on basis of observations. [10 Marks]

b) Find the Hubble constant, $H(t)$, as a function of time in Newtonian cosmological model with $k = 0$. [10 Marks]

c) Find the age of the Universe predicted by the Newtonian theory in this case and express your results in terms of observable cosmological parameters. [10 Marks]