

Problem set 3, Solutions

25 Q1. From Friedmann eq. ~~⊗~~

$$H^2 = \frac{8\pi G \rho}{3} - \frac{kc^2}{a^2} \quad \text{and definition of } \Omega$$

$$\Omega = \frac{8\pi G \rho}{3H^2} \Rightarrow H^2 = \frac{8\pi G \rho}{3\Omega} \quad \text{and}$$

$$\frac{8\pi G \rho}{3\Omega} = \frac{8\pi G \rho}{3} - \frac{kc^2}{a^2}$$

$$\frac{8\pi G \rho}{3} (1 - \Omega^{-1}) = \frac{kc^2}{a^2} \Rightarrow \rho a^2 (1 - \Omega^{-1}) = \frac{3kc^2}{8\pi G} = \text{const} \quad (10)$$

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$$\text{From (10)} \Rightarrow \rho_{eq} a_{eq}^2 (1 - \Omega^{-1})_{eq} = \rho_0 a_0^2 (1 - \Omega^{-1})_0 \quad (11)$$

If $\rho_{rad} \ll \rho_{matter} \Rightarrow$

$$\rho_{eq} = \rho_0 \left(\frac{a_0}{a_{eq}} \right)^3 \quad (12)$$

From (11) and (12) ~~⊗~~ and $\Omega_0 = 0.3$

$$\rho_0 a_0^3 a_{eq}^{-3} a_{eq}^2 (1 - \Omega^{-1})_{eq} = \rho_0 a_0^2 \left(1 - \frac{10}{3} \right)$$

$$(1 - \Omega^{-1})_{eq} = -\frac{a_{eq}}{a_0} \cdot \frac{7}{3} = -\frac{7}{3(1+z_{eq})}$$

$$1 - \Omega_{eq}^{-1} = -\frac{7}{3(1+z_{eq})} \quad \Omega_{eq}^{-1} = 1 + \frac{7}{3(1+z_{eq})}$$

$$\Omega_{eq} = \left[1 + \frac{7}{3(1+z_{eq})} \right]^{-1} \approx 1 - \frac{7}{3 \cdot 6000} \approx 1 - 4 \cdot 10^{-4}$$

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This means that at the epoch of ~~equip~~ matter-radiation equality the universe was surprisingly close to flat.

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25 Q2. From Friedmann eq. with $k=0$ and radiation domination we have

$$H^2 = \frac{8\pi G \rho_0 a_0^4}{3 a^4} = H_0^2 \frac{a_0^4}{a^4}$$

From the general expression (5.27) [Lecture 5], we have

$$t_0 = \int_0^{a_0} \frac{da}{aH} = \int_0^{a_0} \frac{da a^2}{a H_0 a_0^2} = \frac{1}{H_0 a_0^2} \int_0^{a_0} a da =$$

$$= \frac{1}{2H_0 a_0^2} a_0^2 = \frac{1}{2H_0} = \frac{1 \text{ Mpc}}{2.72 \text{ km s}^{-1}}$$

$$= \frac{3 \cdot 10^6 \cdot 3 \cdot 10^8 \text{ M} \cdot \text{s}^{-1} \cdot 1 \text{ year}^2}{1.44 \cdot 10^5 \text{ M} \cdot \text{s}^{-1}} \approx 6.10 \text{ year}^2 \cdot 12 = 66 \text{ year}$$

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Repeating the same calculation for a pressureless flat universe, we have

$$H^2 = \frac{8\pi G \rho_0 a_0^3}{3 a^3} = H_0^2 \frac{a_0^3}{a^3}$$

$$t_0 = \int_0^{a_0} \frac{da}{aH} = \int_0^{a_0} \frac{da a^{3/2}}{a H_0 a_0^{3/2}} = \frac{1}{H_0 a_0^{3/2}} \int_0^{a_0} da a^{1/2} =$$

$$= \frac{2}{3H_0 a_0^{3/2}} a_0^{3/2} = \frac{2}{3H_0}, \text{ hence } \frac{t_0}{\tilde{t}_0} = \frac{3}{4}$$

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25 Q3. See solution in lecture 5 (3)
(eqs 5.28-5.40) and demonstrate
that you do understand what you
are doing.

25 Q4. See solution in lecture 5
(eqs. 5.41-5.46). the same task as
in Q3.