

H. Solutions to course work 4

Q1. 25 Marks

Using a locally-inertial coordinate system prove that the Riemann tensor has the following symmetry properties:

a) Permutation within a pair of indices

$$R_{iklm} = -R_{kilm} = -R_{ikml}.$$

Solution:

$$R_{iklm}^i = \Gamma_{km,l}^i - \Gamma_{kl,m}^i + \Gamma_{nl}^i \Gamma_{km}^n - \Gamma_{nm}^i \Gamma_{kl}^n, \quad (\text{H.1})$$

and

$$R_{iklm} = g_{ip} R_{iklm}^p = g_{ip} \left(\Gamma_{km,l}^p - \Gamma_{kl,m}^p + \Gamma_{nl}^p \Gamma_{km}^n - \Gamma_{nm}^p \Gamma_{kl}^n \right). \quad (\text{H.2})$$

To prove any tensor identity it is enough to show that this identity is true just in the local Galilean frame of reference. In the local Galilean frame

$$\Gamma_{nl}^i = 0, \quad \text{while } \Gamma_{km,l}^p \neq 0,$$

hence

$$R_{iklm} = \eta_{ip} \left(\Gamma_{km,l}^p - \Gamma_{kl,m}^p \right). \quad (\text{H.3})$$

Taking into account that

$$\Gamma_{km,l}^p = \frac{\eta^{pn}}{2} (g_{kn,m} + g_{nm,k} - g_{km,n}), \quad (\text{H.4})$$

we have

$$\begin{aligned} R_{iklm} &= \frac{1}{2} \eta_{ip} [\eta^{pv} (g_{kv,m} + g_{mv,k} - g_{km,v})_l - \eta^{pv} (g_{kv,l} + g_{lv,k} - g_{kl,v})_m] = \\ &= \frac{1}{2} \delta_i^v (g_{kv,m,l} + g_{mv,k,l} - g_{km,v,l} - g_{kv,l,m} - g_{lv,k,m} + g_{kl,v,m}). \end{aligned}$$

Taking into account that

$$g_{kv,m,l} = g_{kv,l,m} \quad \text{and} \quad g_{ik} = g_{ki},$$

we have

$$R_{iklm} = \frac{1}{2} (g_{im,k,l} + g_{kl,i,m} - g_{il,k,m} - g_{km,i,l}). \quad (\text{H.5})$$

To calculate R_{kilm} let us produce the following permutation in (H.5): $i \leftrightarrow k$; then

$$R_{kilm} = \frac{1}{2} (g_{km,i,l} + g_{il,k,m} - g_{kl,i,m} - g_{im,k,l}) = -\frac{1}{2} (g_{im,k,l} + g_{kl,i,m} - g_{il,k,m} - g_{km,i,l}) = -R_{iklm}.$$

To calculate R_{ikml} let us produce the following permutation in (H.5): $l \leftrightarrow m$; then

$$R_{ikml} = \frac{1}{2} (g_{il,k,m} + g_{km,i,l} - g_{im,k,l} - g_{kl,i,m}) = -\frac{1}{2} (g_{im,k,l} + g_{kl,i,m} - g_{il,k,m} - g_{km,i,l}) = -R_{iklm}.$$

b) Permutation of pairs

$$R_{iklm} = R_{lmik}.$$

Solution:Producing the permutation $(ik) \leftrightarrow (lm)$ in (H.5), we have

$$R_{lmik} = \frac{1}{2}(g_{lk,m,i} + g_{mi,l,k} - g_{li,m,k} - g_{mk,l,i}) = \frac{1}{2}(g_{im,k,l} + g_{kl,i,m} - g_{il,k,m} - g_{km,i,l}) = R_{iklm}.$$

Q2. 25 Marks

a) Prove that

$$R_{iklm} + R_{imkl} + R_{ilmk} = 0.$$

Solution:With the help of circling permutation

$$\begin{array}{c} \rightarrow l \\ k \quad \downarrow \\ \leftarrow m \end{array}$$

we have

$$\begin{aligned} & R_{iklm} + R_{imkl} + R_{ilmk} = \\ &= \frac{1}{2}[(g_{im,k,l} + g_{kl,i,m} - g_{il,k,m} - g_{km,i,l}) + (g_{ik,l,m} + g_{lm,i,k} - g_{im,l,k} - g_{lk,i,m}) + (g_{il,m,k} + g_{mk,i,l} - g_{ik,m,l} - g_{ml,i,k})] = \\ &= \frac{1}{2}[(g_{im,k,l} - g_{im,l,k}) + (g_{kl,i,m} - g_{lk,i,m}) - (g_{il,k,m} - g_{il,m,k}) - (g_{km,i,l} - g_{mk,i,l}) + (g_{ik,l,m} - g_{ik,m,l}) + (g_{lm,i,k} - g_{ml,i,k})] = 0. \end{aligned}$$

b) Using a locally-inertial coordinate system prove the Bianchi identity:

$$R_{ikl;m}^n + R_{imk;l}^n + R_{ilm;k}^n = 0.$$

Solution:In a locally-inertial frame of reference

$$R_{iklm}^i = \Gamma_{km,l}^i - \Gamma_{kl,m}^i$$

and

$$R_{ikl;m}^n + R_{imk;l}^n + R_{ilm;k}^n = R_{ikl,m}^n + R_{imk,l}^n + R_{ilm,k}^n.$$

With the same circling permutation of indices as before

$$\begin{array}{c} \rightarrow l \\ k \quad \downarrow \\ \leftarrow m \end{array}$$

we obtain

$$\begin{aligned} & R_{ikl;m}^n + R_{imk;l}^n + R_{ilm;k}^n = \\ &= (\Gamma_{il,k}^n - \Gamma_{ik,l}^n)_{,m} + (\Gamma_{im,l}^n - \Gamma_{il,m}^n)_{,k} + (\Gamma_{ik,m}^n - \Gamma_{im,k}^n)_{,l} = \\ &= \Gamma_{il,k,m}^n - \Gamma_{ik,l,m}^n + \Gamma_{im,l,k}^n - \Gamma_{il,m,k}^n + \Gamma_{ik,m,l}^n - \Gamma_{im,k,l}^n = \\ &= (\Gamma_{il,k,m}^n - \Gamma_{il,m,k}^n) - (\Gamma_{ik,l,m}^n - \Gamma_{ik,m,l}^n) + (\Gamma_{im,l,k}^n - \Gamma_{im,k,l}^n) = 0. \end{aligned}$$

Q3. 25 Marks

a) Give the definition of the Ricci tensor R_{ik} and prove that

$$R_{ik} = \Gamma_{ik,l}^l - \Gamma_{il,k}^l + \Gamma_{ik}^l \Gamma_{lm}^m - \Gamma_{il}^m \Gamma_{km}^l.$$

Solution:

$$R_{ik} = g^{lm} R_{limk} = R_{ilk}^l.$$

Producing summation in the expression (H.1) for the Riemann tensor $i = l$ and replacing k by i and m by k we have

$$\begin{aligned} R_{ik} &= R_{ilk}^l = \Gamma_{ik,l}^l - \Gamma_{il,k}^l + \Gamma_{nl}^l \Gamma_{ik}^n - \Gamma_{nk}^l \Gamma_{il}^n = \\ &= \Gamma_{ik,l}^l - \Gamma_{il,k}^l + \Gamma_{ik}^l \Gamma_{lm}^m - \Gamma_{il}^m \Gamma_{km}^l. \end{aligned}$$

b) Using the Bianchi identity, prove that the Ricci tensor and the scalar curvature $R = g^{ik} R_{ik}$ satisfy the following identity:

$$R_{m;l}^l = \frac{1}{2} \frac{\partial R}{\partial x^m}.$$

Solution: After contracting the Bianchi identity

$$R_{klm;n}^i + R_{knl;m}^i + R_{kmn;l}^i = 0$$

over indices i and n (taking summation $i = n$) we obtain

$$R_{klm;i}^i + R_{kil;m}^i + R_{kmi;l}^i = 0.$$

According to the definition of Ricci tensor

$$R_{kil}^i = R_{kl},$$

the second term can be rewritten as

$$R_{kil;m}^i = R_{kl;m}.$$

Taking into account that the Riemann tensor is antisymmetric with respect permutations of indices within the same pair

$$R_{kmi}^i = -R_{kim}^i = -R_{km},$$

the third term can be rewritten as

$$R_{kmi;l}^i = -R_{km;l}.$$

The first term can be rewritten as

$$R_{klm;i}^i = g^{ip} R_{pklm;i},$$

then taking mentioned above permutation twice we can rewrite the first term as

$$R_{klm;i}^i = g^{ip} R_{pklm;i} = -g^{ip} R_{kplm;i} = g^{ip} R_{kpml;i}.$$

After all these manipulations we have

$$g^{ip} R_{kpml;i} + R_{kl;m} - R_{km;l} = 0.$$

Then multiplying by g^{km} and taking into account that all covariant derivatives of the metric tensor are equal to zero, we have

$$g^{km} g^{ip} R_{kpml;i} + g^{km} R_{kl;m} - g^{km} R_{km;l} = (g^{km} g^{ip} R_{kpml})_{;i} + (g^{km} R_{kl})_{;m} - (g^{km} R_{km})_{;l} = 0.$$

In the first term expression in brackets can be simplified as

$$g^{km}g^{ip}R_{kpml} = g^{ip}R_{pl} = R_l^i.$$

In the second term expression in brackets can be simplified as

$$g^{km}R_{kl} = R_l^m.$$

According to the definition of scalar curvature

$$R = g^{km}R_{km},$$

the third term can be simplified as

$$(g^{km}R_{km})_{;l} = R_{;l} = R_{,l}.$$

Thus

$$R_{l;i}^i + R_{l;m}^m - R_{,l} = 0,$$

replacing in the second term index of summation m by i we finally obtain

$$2R_{l;i}^i - R_{,l} = 0,$$

finally

$$R_{l;i}^i - \frac{1}{2}R_{,l} = 0. \tag{H.6}$$

Q4. 25 Marks

a) Starting from the Einstein equations in the form

$$R_{ik} = \frac{8\pi G}{c^4} \left(T_{ik} - \frac{1}{2}g_{ik}T \right),$$

prove that

$$T_k^i = \frac{c^4}{8\pi G} \left(R_k^i - \frac{1}{2}\delta_k^i R \right),$$

where δ_k^i is the unit diagonal four-tensor.

Solution: Let us first rewrite the EFEs in the mixed form

$$R_n^i = g^{ik}R_{kn} = \frac{8\pi G}{c^4} \left(g^{ik}T_{kn} - \frac{1}{2}g^{ik}g_{kn}T \right) = \frac{8\pi G}{c^4} \left(T_n^i - \frac{1}{2}\delta_n^i T \right).$$

Then by summation $i = n$ we obtain that

$$R = R_n^n = \frac{8\pi G}{c^4} \left(T_n^n - \frac{1}{2}\delta_n^n T \right) = \frac{8\pi G}{c^4} \left(T - \frac{1}{2}4T \right) = \frac{8\pi G}{c^4} (T - 2T) = -\frac{8\pi G}{c^4} T,$$

hence

$$T_k^i = \frac{c^4}{8\pi G} R_k^i + \frac{1}{2}\delta_k^i T = \frac{c^4}{8\pi G} \left(R_k^i - \frac{1}{2}\delta_k^i R \right).$$

b) What can you say about the nature of gravitational field, for which $R_{ik} = 0$, while R_{ikln} is not equal to zero?

Solution: $R_{ik} = 0$ implies that $T_{ik} = 0$, i.e. the space-time is empty. $R_{ikln} \neq 0$ means that the space-time is not flat. [Thus we have free gravitational field in empty space-time which is gravitational wave.]

c) Prove that the energy-momentum tensor of matter T_k^i satisfies the conservation law $T_{k;i}^k = 0$.

Solution: Taking covariant divergence of the EFEs in mixed form we obtain

$$R_{k;i}^i - \frac{1}{2}\delta_k^i R_{,i} = \frac{8\pi G}{c^4} T_{k;i}^i. \tag{H.7}$$

As follows from (H.6)(see Q3b) the LHS of (H.7) is equal to zero, hence

$$T_{k;i}^i = 0.$$